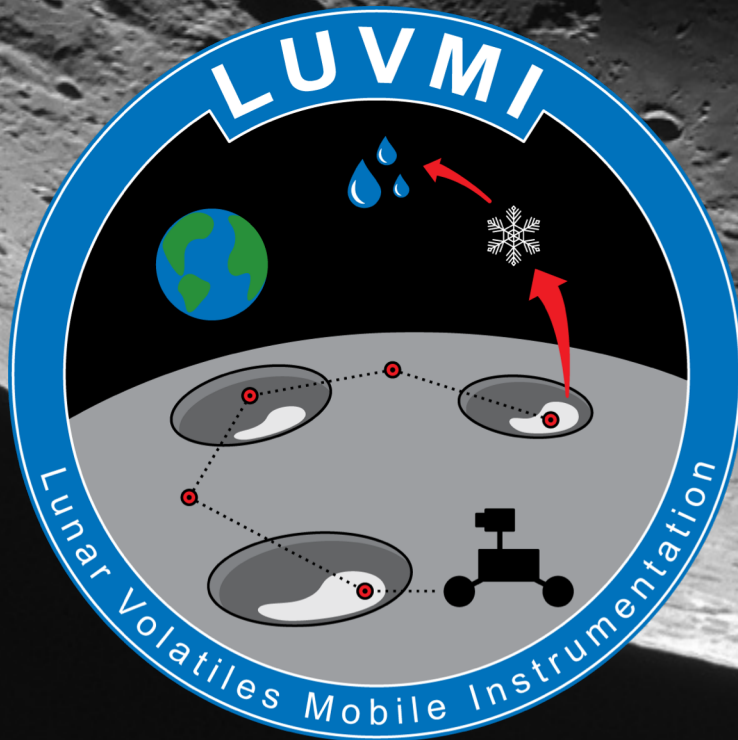


LUVMI: low-mass, low-footprint, payload and robotic system for the sampling of volatiles at the Lunar poles

Space Resources Roundtable, Golden, CO



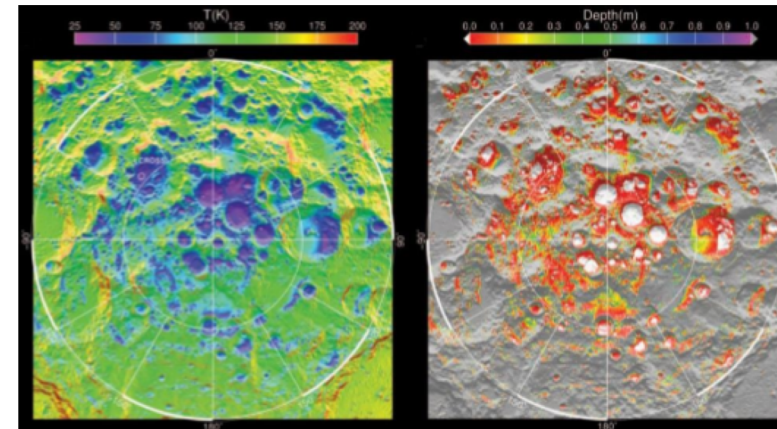
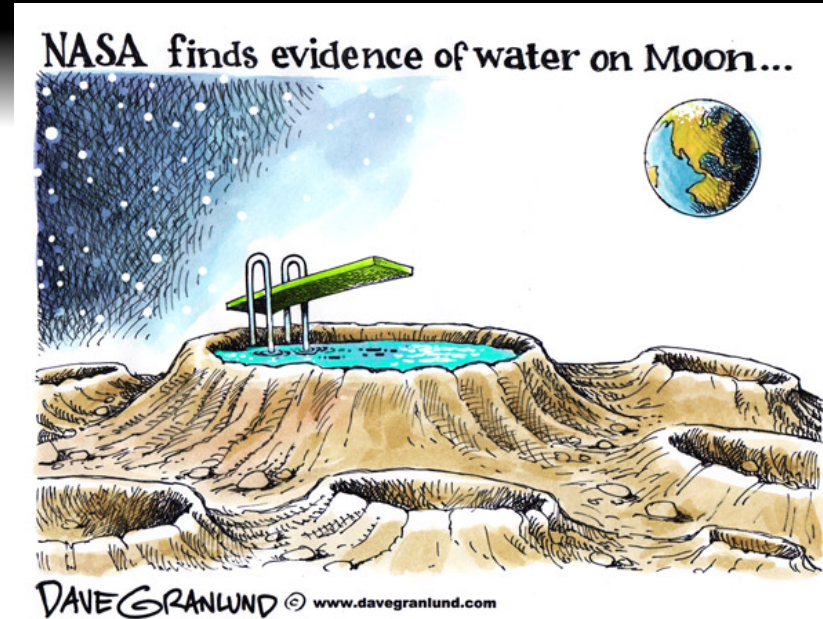
*Diego A. Urbina, Team Lead Future Projects and Exploration,
Space Applications Services, Belgium*

**Jeremi Gancet, Karsten Kullack, Enrico Ceglia Hemanth K.
Madakashira, Joseph Salini, Shashank Govindaraj, Leonardo
Surdo, Richard Aked, Simon Sheridan, Craig Pitcher, Simeon
Barber, Janos Biswas, Philipp Reiss, Joseph Rushton, Neil
Murray, Anthony Evangora, Lutz Richter, Diana Dobrea,
Mattia Reganaz**

Context & motivation



- Lunar volatiles – untapped resource to kick start the space economy
- H, H₂O, N, CO₂, others ... but where, how much?
- Prospecting missions planned but heavy-> few, far in between
- Use low-mass robotic payload



D. Paige, M. A. (2010). Diviner Lunar Radiometer Observations of Cold Traps in the Moon's South Polar Region. *Science*, 479-482.

- LUVMI: LUnar Volatiles Mobile Instrumentation
- Project H2020 (COMPET-5-2016: scientific instrumentation)
- 24 months

Objectives



- Objectives:
 - Determine variability of volatile distribution (VSAT obj. 1)
 - *Identification of chemical phase of volatile elements* (VSAT obj. 2)
 - Analysis of physical & chemical behavior of lunar soil with T° (VSAT obj. 3)
 - Determine geotechnical properties (VSAT obj. 4)
 - Current volatile flux (VSAT obj. 5)
- A small mobile payload addresses 4 objectives (+1 partially)

	Static lander with limited subsurface access	Static Lander with drill	Rover		
			Surface sampling only	Limited subsurface access (<20 cm)	drill (>1 m)
Variability of volatile distribution	1	2			
Chemical Phase of Volatile Elements	3		4	5	
Chemical and physical behavior of polar soil with temperature			6		
Geotechnical properties			7		
Current volatile flux					

Credits LEAG VSAT

Scientific objectives



- Scientific objectives:
 - Detect/quantify abundance of loosely bound volatiles
 - Identify chemical phase of detected volatiles
 - Investigate migration of volatiles over lunar surface
 - Determine engineering properties of regolith in polar regions.
- Expected detection:
 - Water
 - Volatiles observed during LCROSS

Mission Baseline

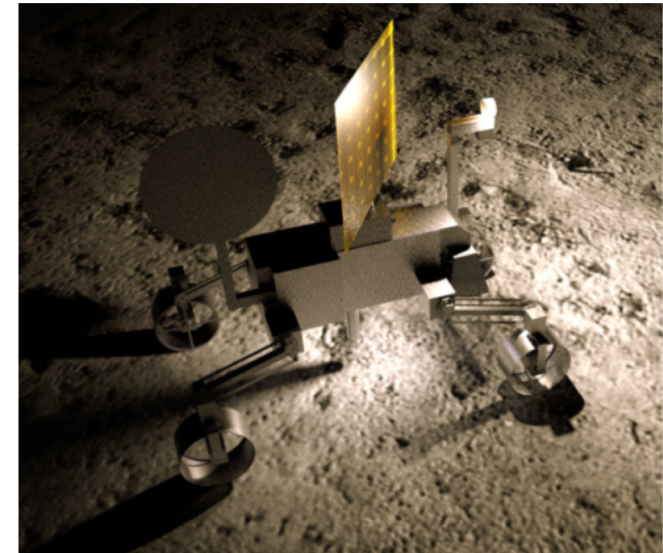
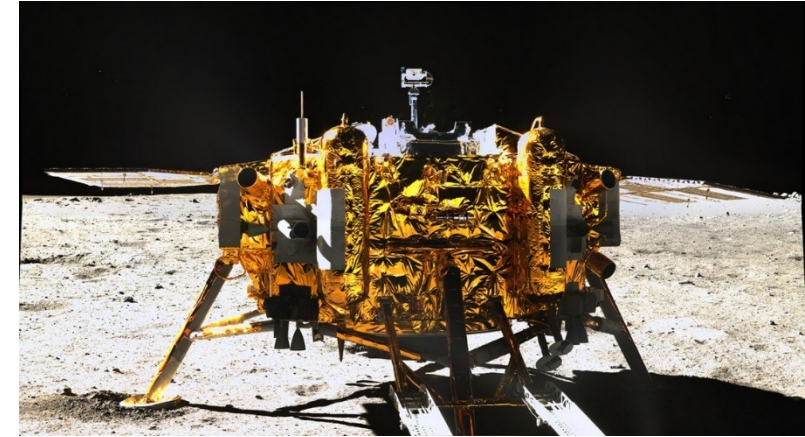


- Landing site:
 - 1km distance to PSR
 - >80 deg latitude
 - 40% solar illumination [avg. time]
 - 40% Earth visibility [avg. time]
 - No major obstacle in 250m radius
- Mission
 - 14 days
 - Increasing order of risk
 - Hibernation possible for 48h
- Sampling locations:
 - Vicinity of landing site
 - Regular illumination sites
 - Non-permanent shadow region
 - Night/day termination event
 - Vicinity of PSR
 - PSR

Mission Baseline



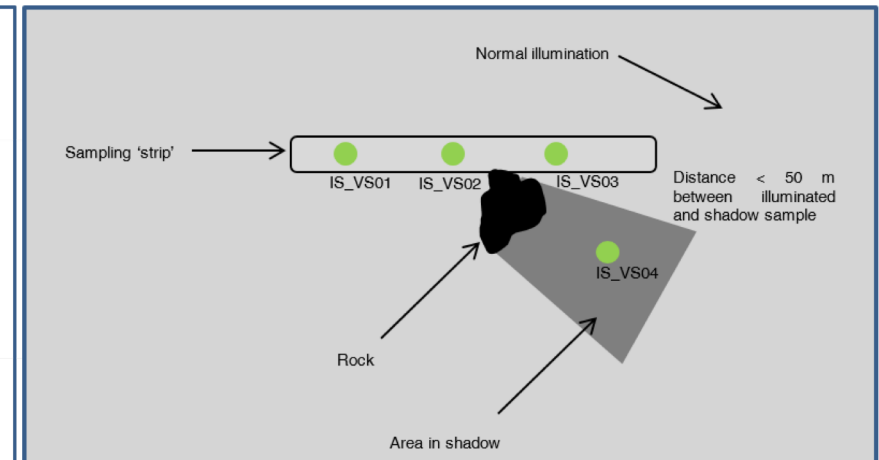
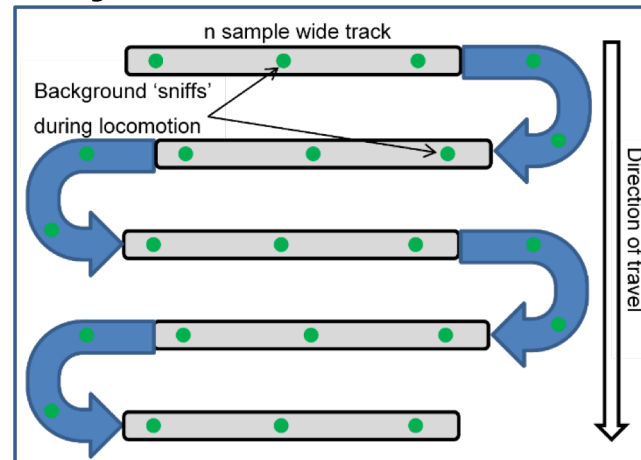
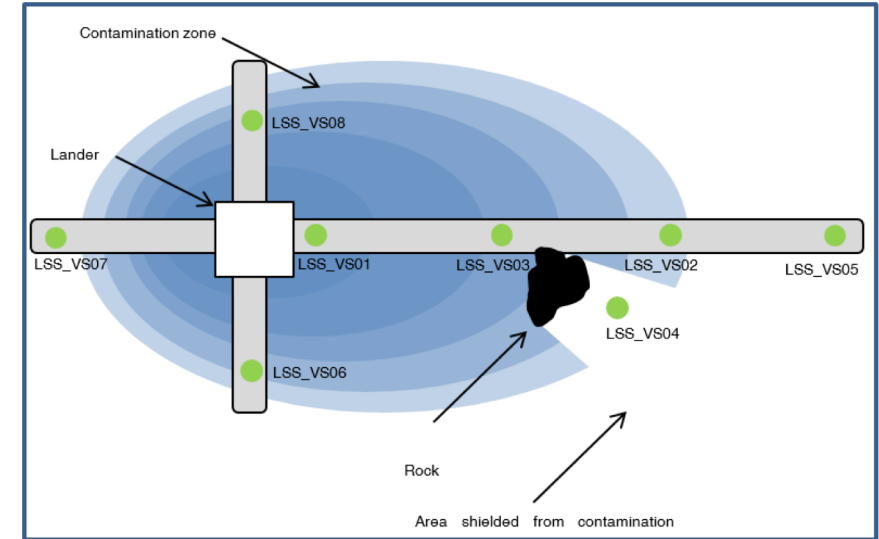
- Flight: Design LUVMI to accommodate most landers
 - Piggy back(Change'e, Luna-[25,27,28,29], NASA Resource Prospector)
 - Commercial lander (Astrobotic , PTScientists...)
- Design LUVMI to withstand PSR environment
- Independent from Lander



Concept of operation

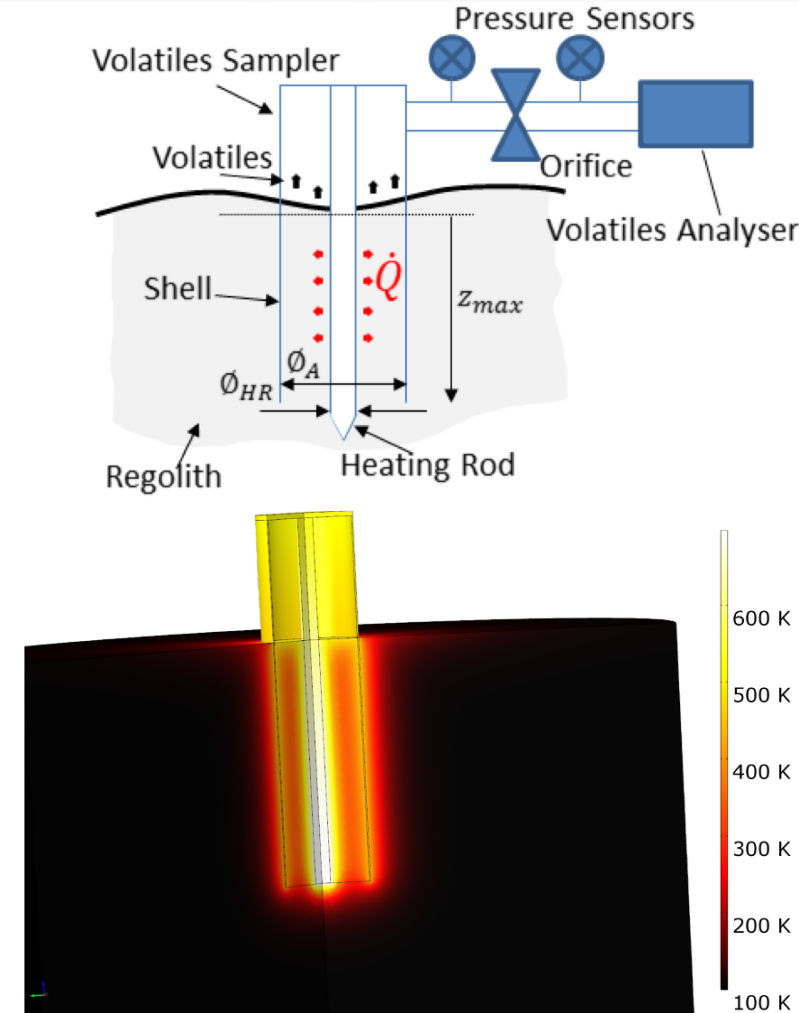
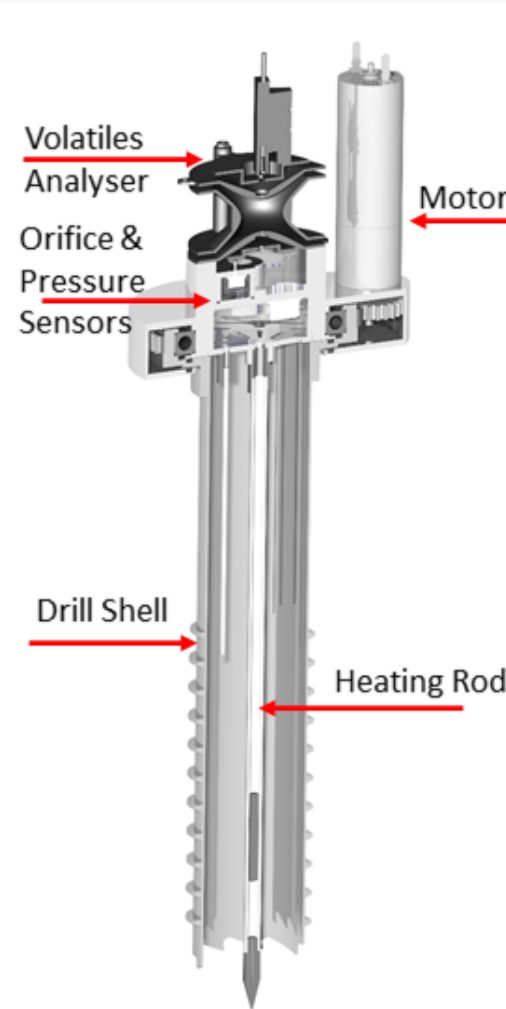


- Post landing-checkout
- Rover egress
- Landing survey
- Wide field area survey
- Terminator passage survey
- PSR survey

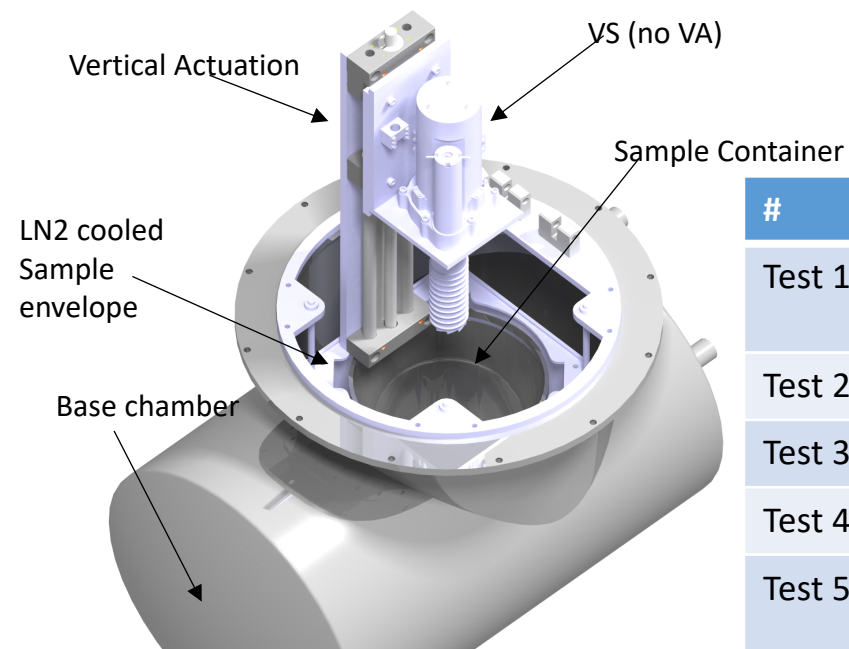


Instruments: Volatile Sampler

- Drill + sample preparation + gas extraction
- Operation:
 - Get rover into drilling position
 - Drill (10 cm deep, possibly up to 20 cm)
 - Heat central rod
 - Extract/analyze trapped volatiles
- Low heat conductivity of Regolith
 - High gradient



Instruments: Volatile Sampler



Volatiles Extraction Test

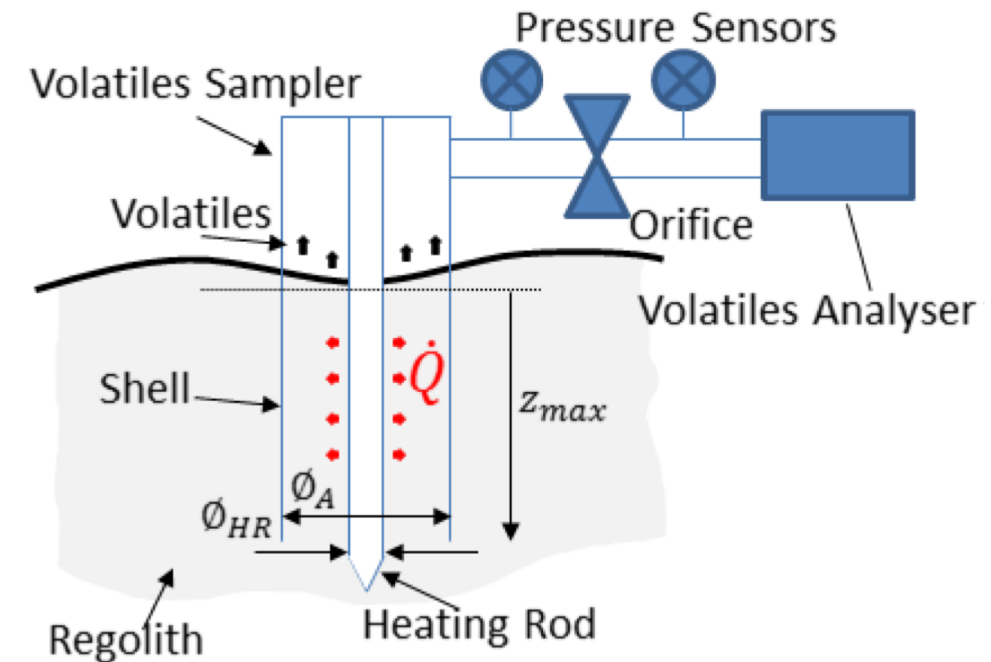
#	Description	Status
Test 1	Mechanism Functionality Test (OHB)	Successful
Test 2	Penetration Test (OHB)	Ongoing
Test 3	Materials Outgassing Test (OHB)	Ongoing
Test 4	Volatiles Extraction Test	In preparation
Test 5	E2E VA/VS Integrated Gas Extraction & Analysis	In preparation
Test 6	Rover Demonstration	-



Penetration test

Instruments: Volatile Analyzer

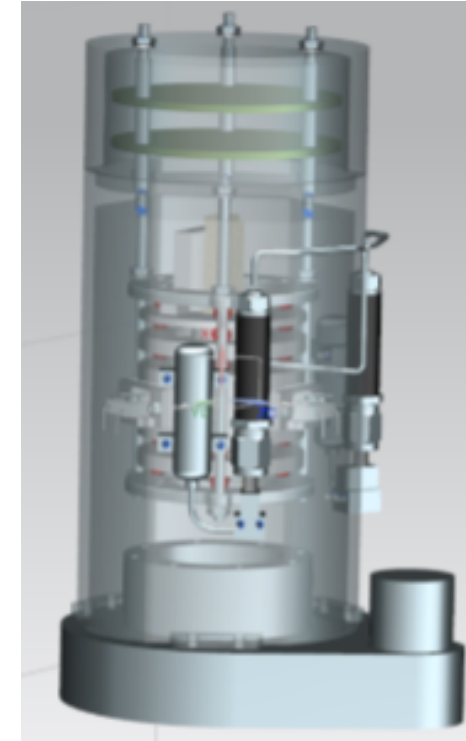
- Placed on top of Volatile sampler
- Ion trap mass spectrometer
- Based on heritage of Rosetta mission and on UK MoonLite mission instrument



Instruments: Volatile Analyzer



- Detect: H_2S , NH_3 , SO_2 , C_2H_4 , CO_2 , CH_3OH , CH_4
- Simple, low mass, compact
- Low detection levels
- Reference gas system, in-situ calibration



LUVMI mass spec (VA)



Rosetta ptolemy mass spec

VA-VS assembly (payload)

- For prototype:
 - Mass VS+VA= 1,8 kg plus margins
 - Mass VS+VA electronics= 1 kg plus margins
 - Max Power during Analysis VS+VA= 26 W



Preliminary integration (End of May 2018)

Instruments: Imager



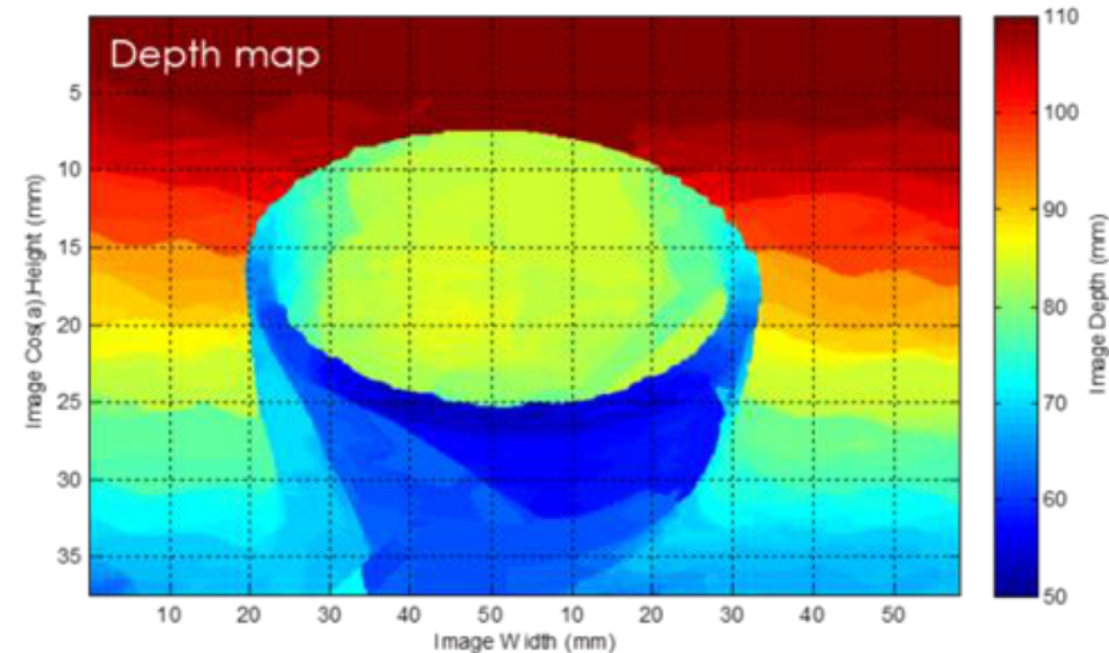
Technology

- Light-field imaging
- Micro-lens array
- Extracts depth map

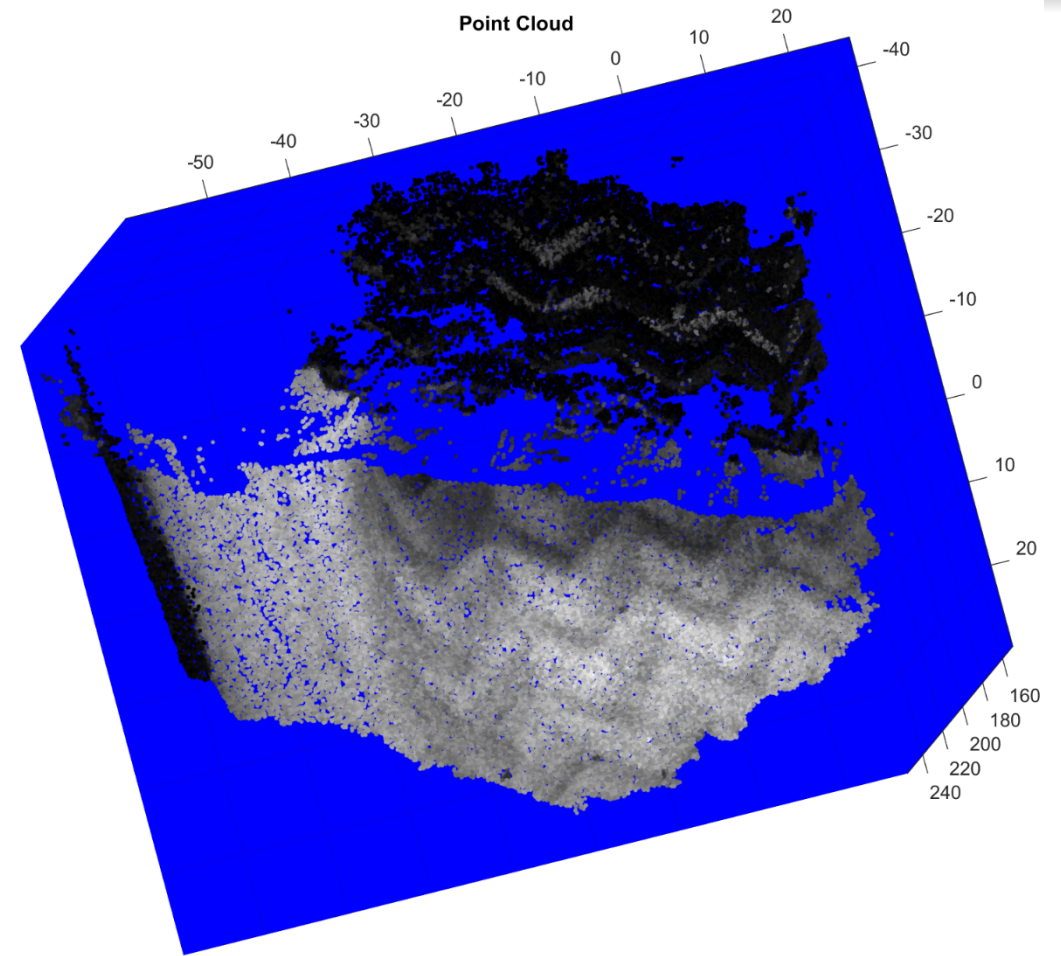
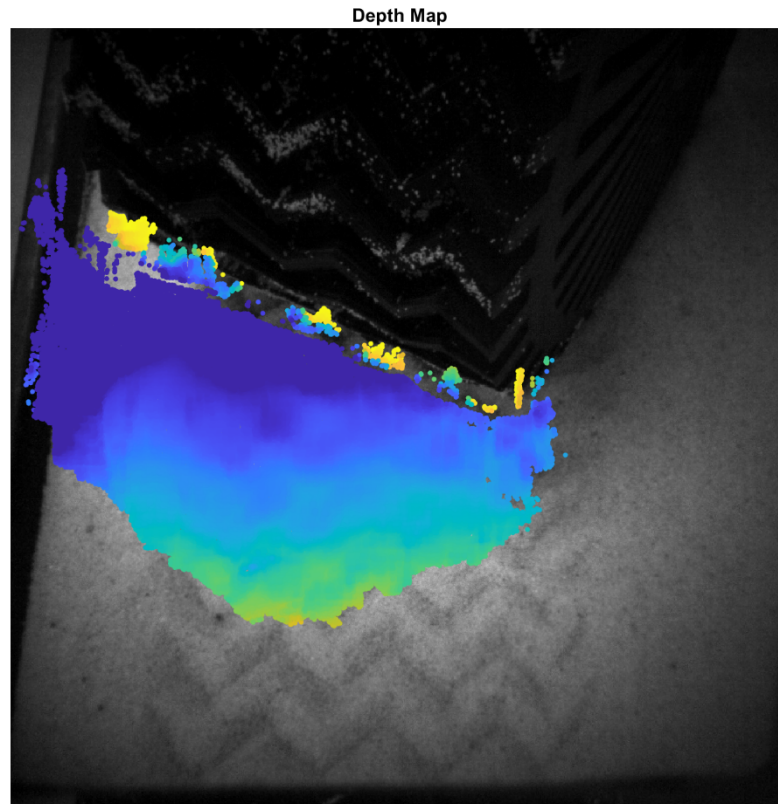


Provides

- Info on rover's environment
- For navigation, teleoperation and science



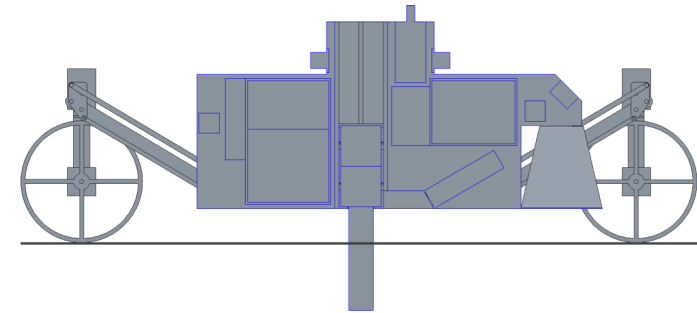
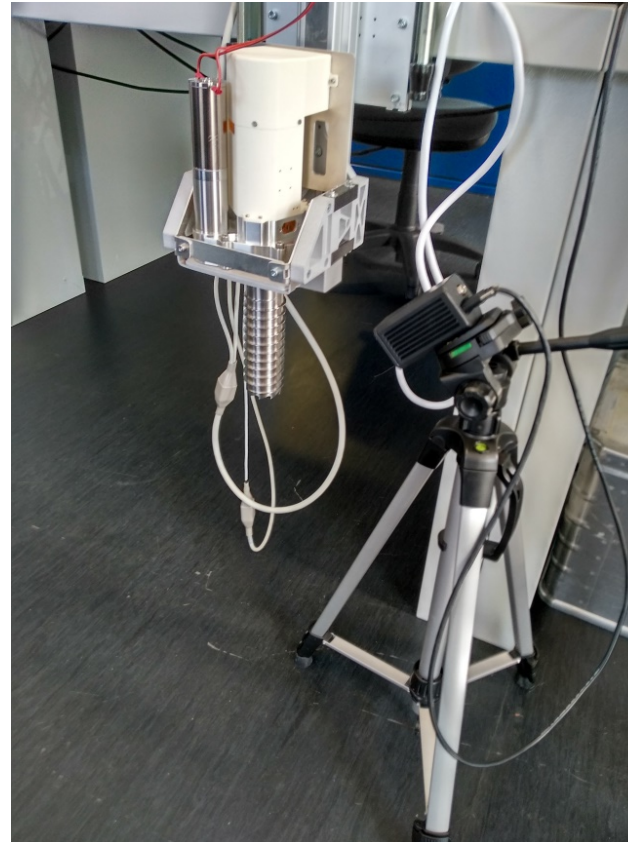
Instruments: Imager



Instruments: Imager

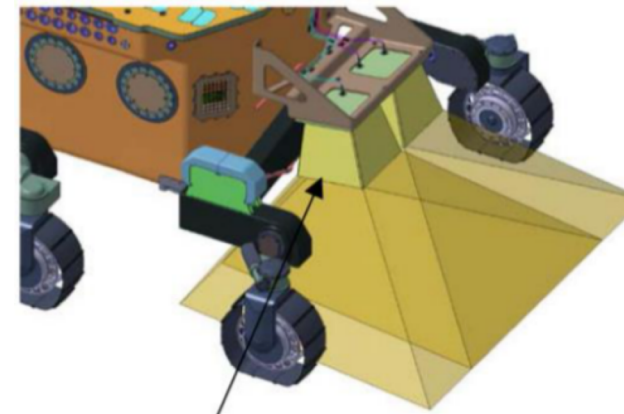


- Imager in mast (navigation)
- Imager under chassis
 - Monitor sampling process
 - Measure geotechnical properties



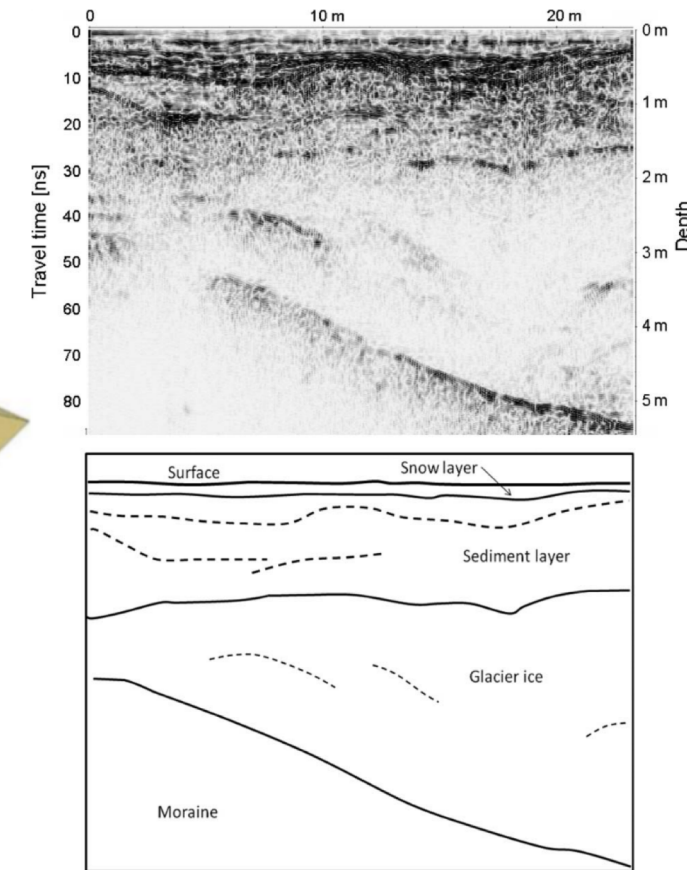
Instruments: Ground penetrating radar

- Possible Secondary payload
 - WISDOM - based on exoMars heritage
- Get subsurface context + correlate with measured volatiles
- Information before drilling
 - Hardness



Antennes WISDOM

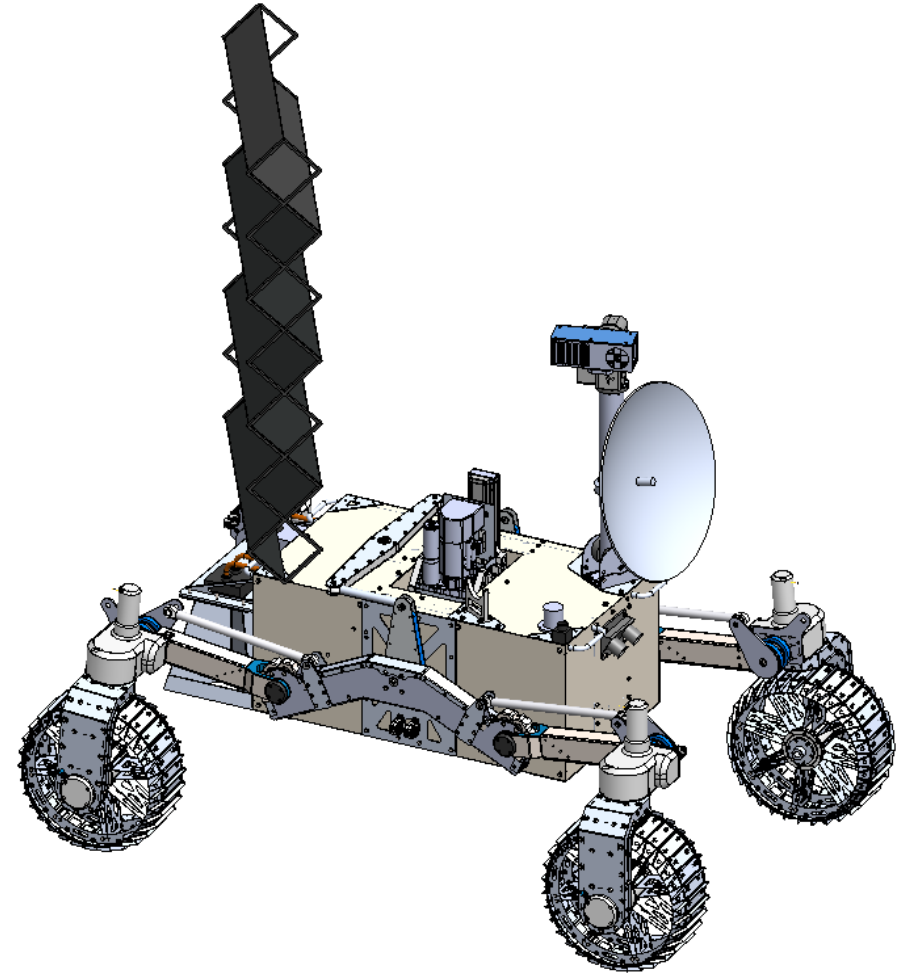
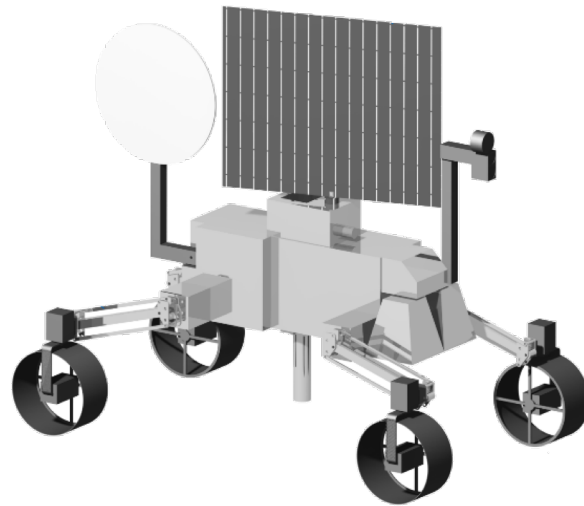
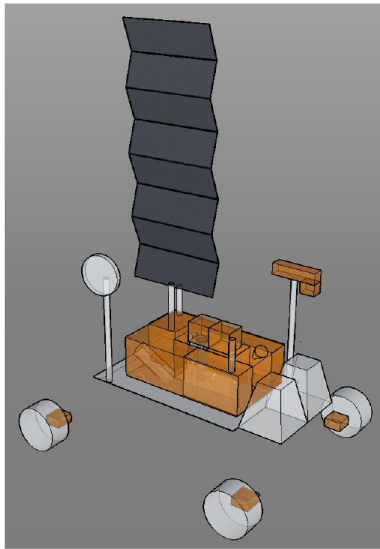
WISDOM on ExoMars rover: Image courtesy of LATMOS



Mobile platform



- Flight Model -> preliminary design
- Ground Model -> prototype
 - Demonstrate critical functions
 - Test in earth-based analogues



Mobile platform



NASA LRP: 1.5m x 1.5m x 2.0m
Weights about 300kg



LUVMI: 1.5m x 1m x 1.6m
(deployed)
Weights about 40 kg



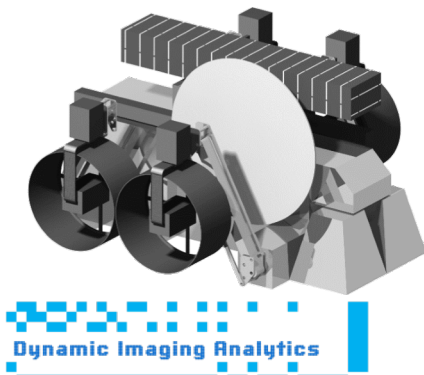
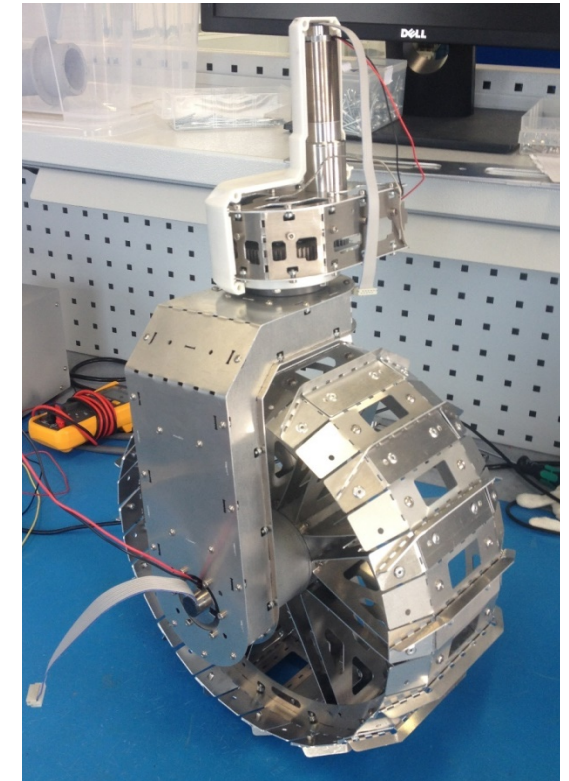
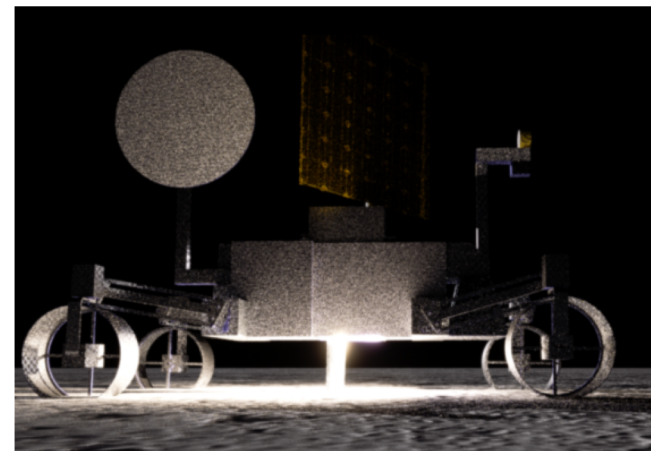
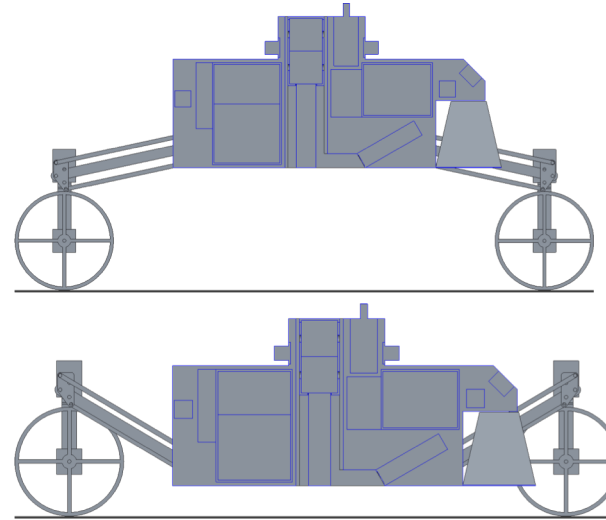
1m x .7m x
.7m (stowed)

- Note LRP has/had many functional differences: sealed manifold, controlled heating, quantitative measurements, water droplet experiment, ilmenite reduction (?), deep access
- LUVMI is conceived as an “LRP LITE” w/ reasonable compromises, achieving a much lighter payload and rover

Mobile platform



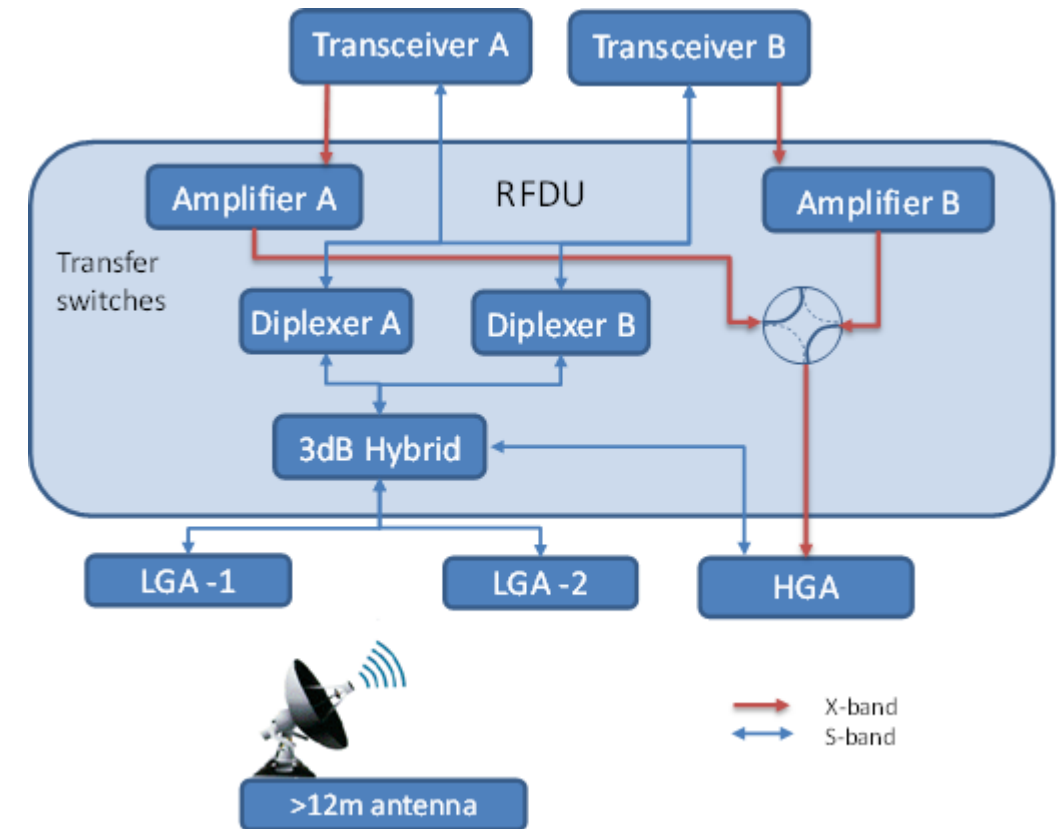
- Active suspension
- Enables drilling operation
- Used for facilitating stowage in lander
- Compact system in lander
- Large wheelbase & front track
- Larger ground clearance



Mobile platform sub-systems



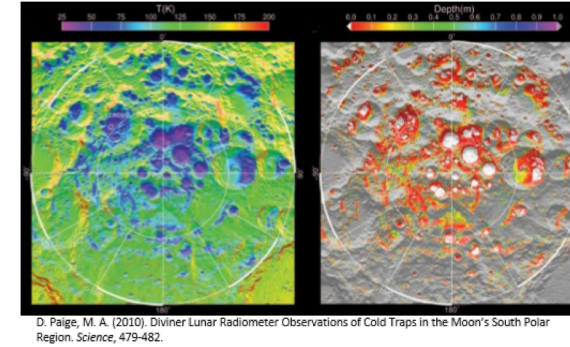
- Communication
 - High/low gain: Direct-To-Earth communication
 - Lander or orbiter comms possible when and if available (not req'd)
- Power
 - Plan to use high-density batteries (Li-S)
 - Will be changed if TRL low by time of building system/ if low temperature battery-less architecture feasible
- Thermal
 - Provide heat to keep system alive
 - in PSR
 - During hibernation



Conclusion & future work



- Sum-up:
 - 40-kg high-mobility rover, survives PSRs. State of the art miniature analyser and sampler, GPR, lightfield imager.
 - Flexible design, can piggyback in agency, commercial missions
 - **Additional payloads welcome**
- On-going and future work
 - Building instrument and mobile platform system
 - Test instruments in vacuum
 - Test integrated system in earth-based analogue
 - **BLTAS: ESA GSP project**
 - Studies battery-less “low-temperature” architecture
 - Using LUVMI as case study



D. Paige, M. A. (2010). Diviner Lunar Radiometer Observations of Cold Traps in the Moon's South Polar Region. Science, 479-482.

